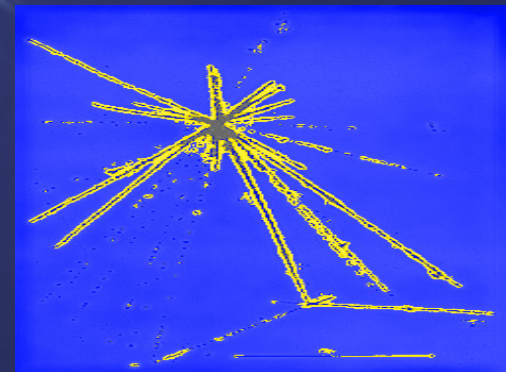


# Supersymmetry in Particle Physics

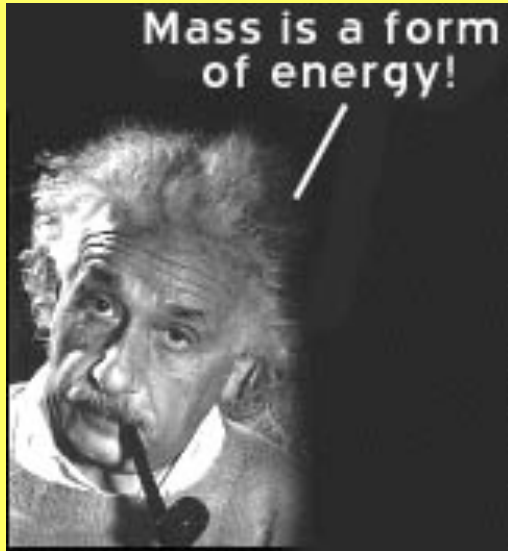
## Outline

1. What is SUSY
2. Basics of SUSY
3. The MSSM
4. Constrained MSSM
5. SUSY Searches
6. SUSY DM

$e^-$   $e^+$

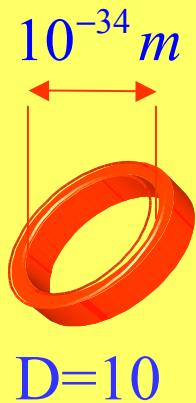
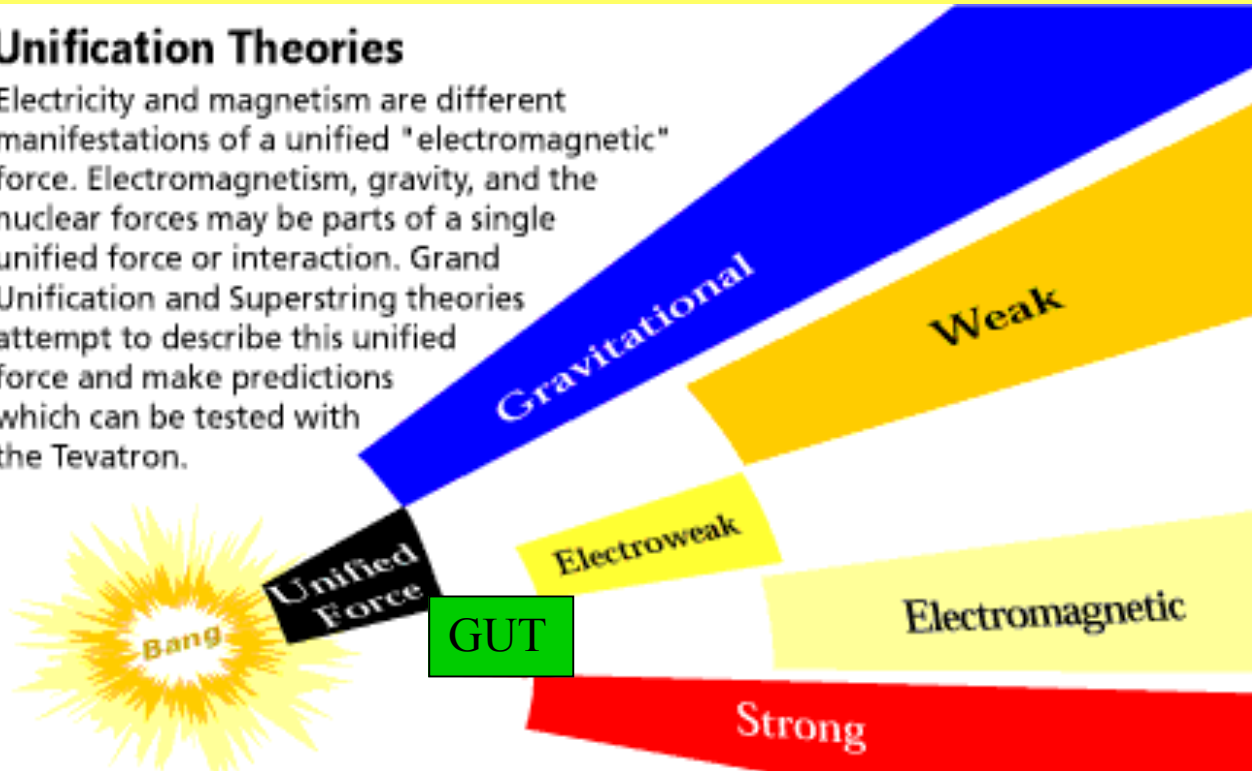


# Unification Paradigm

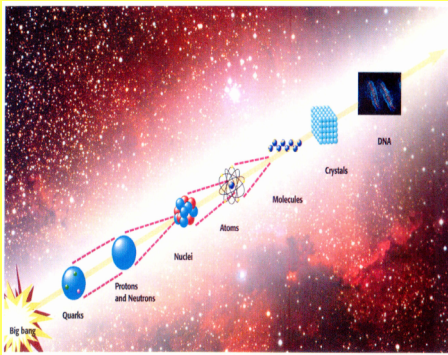


## Unification Theories

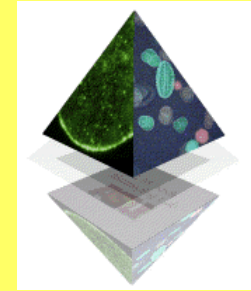
Electricity and magnetism are different manifestations of a unified "electromagnetic" force. Electromagnetism, gravity, and the nuclear forces may be parts of a single unified force or interaction. Grand Unification and Superstring theories attempt to describe this unified force and make predictions which can be tested with the Tevatron.



- Unification of strong, weak and electromagnetic interactions within Grand Unified Theories is the new step in unification of all forces of Nature
- Creation of a unified theory of everything based on string paradigm seems to be possible



# What is SUSY?



- **Supersymmetry** is a boson-fermion symmetry that is aimed to unify all forces in nature (including gravity) within a single framework

$$Q |boson \rangle = |fermion \rangle$$

$$[b, b^\dagger] = 1 \quad \sigma^\mu (\sigma^\nu)_{\alpha\beta} P_\mu$$

First papers in 1971-1972  
 No evidence in particle physics yet

- Modern theories in particle physics are based on supersymmetry, though low energy manifestation of SUSY can be found (?) at modern colliders and in non-accelerator experiments

# Motivation of SUSY in Particle Physics

- **Unification with Gravity**
- Unification of gauge couplings
- Solution of the hierarchy problem  $Q|boson\rangle = |fermion\rangle, Q|fermion\rangle = |boson\rangle$
- Dark matter in the Universe
- Superstrings  $spin\ 2 \rightarrow spin\ 3/2 \rightarrow spin\ 1 \rightarrow spin\ 1/2 \rightarrow spin\ 0$

Unification of matter (fermions) with forces (bosons) naturally arises

from an attempt to unify gravity with the other interactions

$$\{Q_\alpha^i, \bar{Q}_\beta^j\} = 2\delta^{ij}(\sigma^\mu)_{\alpha\beta}P_\mu \Rightarrow \{\delta_\varepsilon, \bar{\delta}_{\bar{\varepsilon}}\} = 2(\varepsilon\sigma^\mu\bar{\varepsilon})P_\mu$$

$$\varepsilon = \varepsilon(x) \text{ local coordinate transf.} \Rightarrow \text{(super)gravity}$$

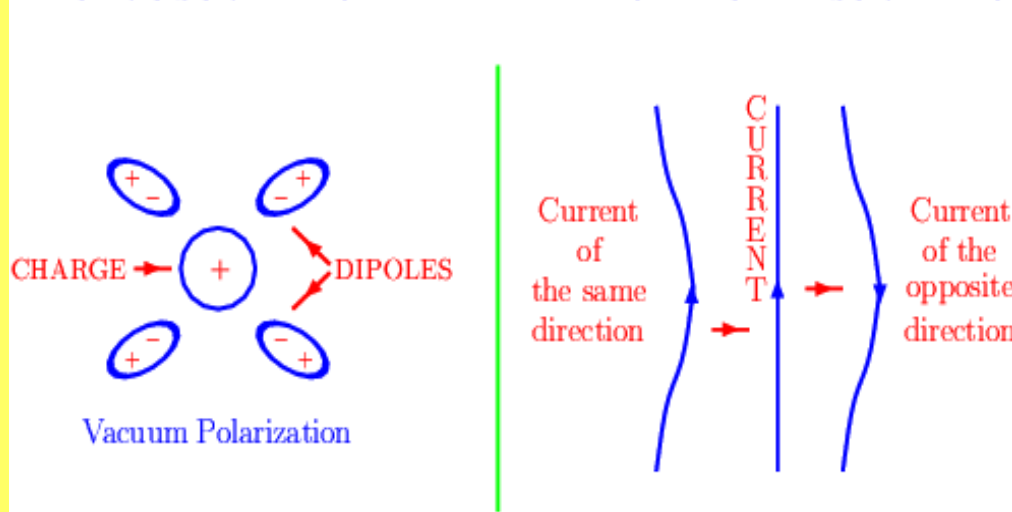
**Local supersymmetry = general relativity !**

# Motivation of SUSY in Particle Physics

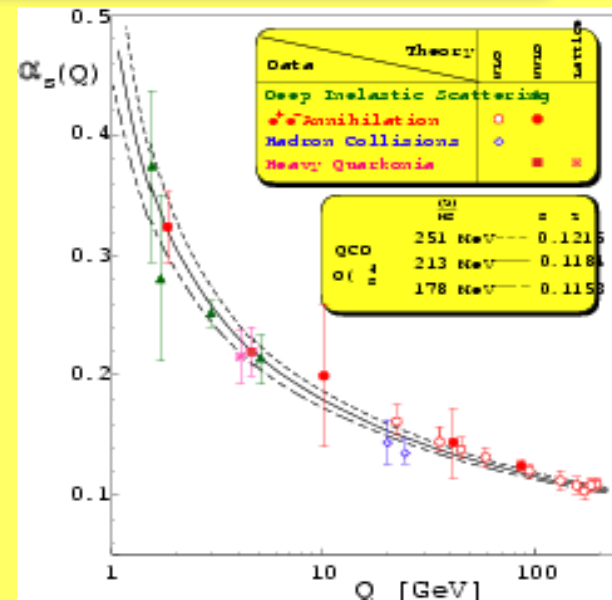
- Unification of gauge couplings

Low Energy			⇒	High Energy
$SU_c(3)$	$SU_L(2)$	$U_Y(1)$	⇒	$G_{GUT}$ (or $G^n + \text{symm}$ )
gluons	$W, Z$	photon	⇒	gauge bosons
quarks	leptons		⇒	fermions
$g_3$	$g_2$	$g_1$	⇒	$g_{GUT}$

ELECTRIC SCREENING



$$\alpha_i = \alpha_i \left( \frac{Q^2}{\Lambda^2} \right) = \alpha_i(\text{distance})$$



Running of the strong coupling

# Motivation of SUSY

RG Equations  $\frac{d\tilde{\alpha}_i}{dt} = b_i \tilde{\alpha}_i^2$ ,  $\tilde{\alpha}_i = \alpha_i / 4\pi = g_i^2 / 16\pi^2$ ,  $t = \log(Q^2 / \mu^2)$

$SM: \quad b_i = \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} = \begin{pmatrix} 0 \\ -22/3 \\ -11 \end{pmatrix} + N_{Fam} \begin{pmatrix} 4/3 \\ 4/3 \\ 4/3 \end{pmatrix} + N_{Higgs} \begin{pmatrix} 1/10 \\ 1/6 \\ 0 \end{pmatrix}$	$MSSM: \quad b_i = \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} = \begin{pmatrix} 0 \\ -6 \\ -9 \end{pmatrix} + N_{Fam} \begin{pmatrix} 2 \\ 2 \\ 2 \end{pmatrix} + N_{Higgs} \begin{pmatrix} 3/10 \\ 1/2 \\ 0 \end{pmatrix}$
--	--

Unification of the Coupling Constants  
in the SM and in the MSSM

Input

$$\alpha^{-1}(M_Z) = 128.978 \pm 0.027$$

$$\sin^2 \theta_{MS} = 0.23146 \pm 0.00017$$

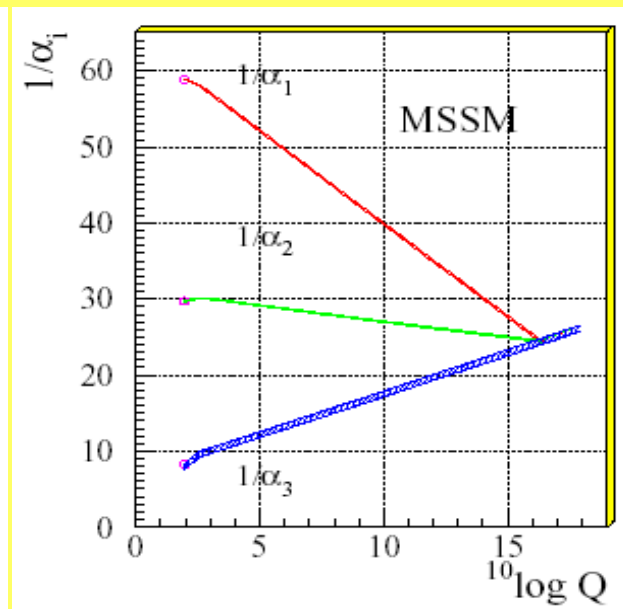
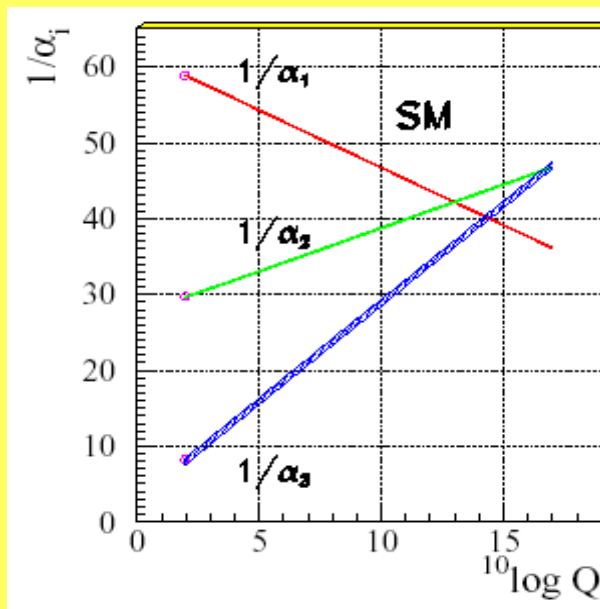
$$\alpha_s(M_Z) = 0.1184 \pm 0.0031$$

Output

$$M_{SUSY} = 10^{3.4 \pm 0.9 \pm 0.4} \text{ GeV}$$

$$M_{GUT} = 10^{15.8 \pm 0.3 \pm 0.1} \text{ GeV}$$

$$\alpha_{GUT}^{-1} = 26.3 \pm 1.9 \pm 1.0$$

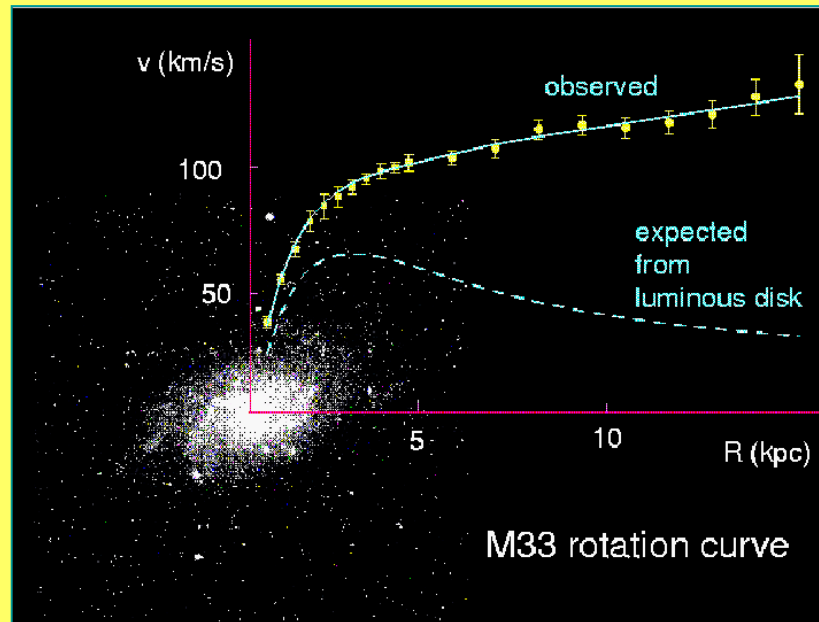


SUSY yields unification!



# Motivation of SUSY

- Dark Matter in the Universe



Spiral galaxies consist of a central bulge and a very thin disc, and surrounded by an approximately spherical halo of dark matter

The flat rotation curves of spiral galaxies provide the most direct evidence for the existence of large amount of the dark matter.



SUSY provides a candidate for the Dark matter – a stable neutral particle



# Cosmological Constraints

New precise cosmological data

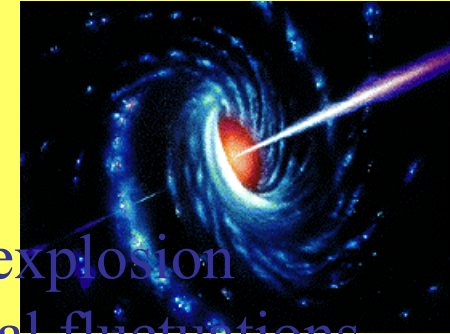
$$\Omega h^2 = 1 \iff \rho = \rho_{crit}$$

$$\Omega_{vacuum} \approx 73\%$$

$$\Omega_{DarkMatter} \approx 23 \pm 4\%$$

$$\Omega_{Baryon} \approx 4\%$$

- Supernova Ia explosion
- CMBR thermal fluctuations  
(recent news from WMAP )



Dark Matter in the Universe:



Hot DM  
(not favoured by  
galaxy formation)

Cold DM  
(rotation curves  
of Galaxies)

SUSY

# Superalgebra

(Super) Algebra

Lorentz Algebra

$$[P_\mu, P_\nu] = 0, [P_\mu, M_{\rho\sigma}] = i(g_{\mu\rho}P_\sigma - g_{\mu\sigma}P_\rho),$$

$$[M_{\mu\nu}, M_{\rho\sigma}] = i(g_{\nu\rho}M_{\mu\sigma} - g_{\nu\sigma}M_{\mu\rho} - g_{\mu\rho}M_{\nu\sigma} + g_{\mu\sigma}M_{\nu\rho}),$$

SUSY Algebra

$$[Q_\alpha^i, P_\mu] = [\bar{Q}_{\dot{\alpha}}^i, P_\mu] = 0,$$

$$[Q_\alpha^i, M_{\mu\nu}] = \frac{1}{2}(\sigma_{\mu\nu})_\alpha^\beta Q_\beta^i, [\bar{Q}_{\dot{\alpha}}^i, M_{\mu\nu}] = -\frac{1}{2}\bar{Q}_{\dot{\beta}}^i (\bar{\sigma}_{\mu\nu})_{\dot{\alpha}}^{\dot{\beta}},$$

$$\{Q_\alpha^i, \bar{Q}_{\dot{\beta}}^j\} = 2\delta^{ij} (\sigma^\mu)_{\alpha\dot{\beta}} P_\mu$$

$$\alpha, \dot{\alpha}, \beta, \dot{\beta} = 1, 2; i, j = 1, 2, \dots, N.$$

The only possible graded Lie algebra that mixes integer and half-integer spins and changes statistics

Superspace

$$x_\mu \rightarrow x_\mu, \theta_\alpha, \bar{\theta}_{\dot{\alpha}}$$

Grassmannian parameters

$$\alpha, \dot{\alpha} = 1, 2$$

$$\vartheta_\alpha^2 = 0, \bar{\vartheta}_{\dot{\alpha}}^2 = 0$$

SUSY Generators

$$Q_\alpha = \frac{\partial}{\partial \vartheta_\alpha} - i\sigma_{\alpha\dot{\alpha}}^\mu \bar{\theta}^{\dot{\alpha}} \partial_\mu$$

$$\bar{Q}_{\dot{\alpha}} = -\frac{\partial}{\partial \bar{\vartheta}_{\dot{\alpha}}} + i\theta_\alpha \sigma_{\alpha\dot{\alpha}}^\mu \partial_\mu$$

$$Q_\alpha^2 = 0, \bar{Q}_{\dot{\alpha}}^2 = 0$$

Supertranslation

$$x_\mu \rightarrow x_\mu + i\theta_\mu \bar{\xi} - i\xi_\mu \bar{\theta},$$

$$\theta \rightarrow \theta + \xi,$$

$$\bar{\theta} \rightarrow \bar{\theta} + \bar{\xi}$$

# Quantum States

Quantum states: Vacuum =  $|E, \lambda\rangle$   $Q|E, \lambda\rangle = 0$

$$[Q_\alpha^i, P_\mu] = [\bar{Q}_\alpha^i, P_\mu] = 0$$

Energy
helicity

State	Expression	# of states
vacuum	$ E, \lambda\rangle$	1
1-particle	$\bar{Q}_i  E, \lambda\rangle =  E, \lambda + 1/2\rangle$	$\binom{N}{1} = N$
2-particle	$\bar{Q}_i \bar{Q}_j  E, \lambda\rangle =  E, \lambda + 1\rangle$	$\binom{N}{2} = \frac{N(N-1)}{2}$
...	...	...
N-particle	$\bar{Q}_1 \bar{Q}_2 \dots \bar{Q}_N  E, \lambda\rangle =  E, \lambda + N/2\rangle$	$\binom{N}{N} = 1$

Total # of states:  $\sum_{k=0}^N \binom{N}{k} = 2^N = 2^{N-1} \text{ bosons} + 2^{N-1} \text{ fermions}$

# SUSY Multiplets

Chiral multiplet  $N = 1, \lambda = 0$

helicity	-1/2	0	1/2
# of states	1	2	1

scalar spinor  
 $(\varphi, \psi)$

Vector multiplet  $N = 1, \lambda = 1/2$

helicity	-1	-1/2	1/2	1
# of states	1	1	1	1

$(\lambda, A_\mu)$   
 spinor vector

Members of a supermultiplet are called **superpartners**

N=4	SUSY YM	helicity	-1	-1/2	0	1/2	1				
	$\lambda = -1$	# of states	1	4	6	4	1				
N=8	SUGRA	helicity	-2	-3/2	-1	-1/2	0	1/2	1	3/2	2
	$\lambda = -2$	# of states	1	8	28	56	70	56	28	8	1

$N \leq 4S$  ← spin

$N \leq 4$  For renormalizable theories (YM)  
 $N \leq 8$  For (super)gravity

# Simplest (N=1) SUSY Multiplets

Bosons and Fermions come in pairs

$$(\varphi, \psi)$$

Spin 0

Scalar

Spin 1/2

Chiral fermion

$$(\lambda, A_\mu)$$

Spin 1/2

Majorana fermion

Spin 1

Vector

$$(\tilde{g}, g)$$

Spin 3/2

Gravitino

Spin 2

Graviton

# SUSY Transformation

N=1 SUSY Chiral supermultiplet:

spin=0

$$\delta_\epsilon A = \sqrt{2}\epsilon\psi,$$

spin=1/2

$$\delta_\epsilon \psi = i\sqrt{2}\sigma^\mu \bar{\epsilon} \partial_\mu A + \sqrt{2}\epsilon F,$$

$$\delta_\epsilon F = i\sqrt{2}\bar{\epsilon}\sigma^\mu \partial_\mu \psi$$

parameter of SUSY transformation  
(spinor)

Auxiliary field

(unphysical d.o.f. needed to close SUSY algebra)

SUSY multiplets



Superfield in Superspace

$$(y = x + i\theta\sigma\bar{\theta})$$

Expansion over  
grassmannian  
parameter

superfield

$$\begin{aligned} \Phi(y, \theta) &= A(y) + \sqrt{2}\theta\psi(y) + \theta\theta F(y) \\ &= A(x) + i\theta\sigma^\mu \bar{\theta} \partial_\mu A(x) + \frac{1}{4}\theta\theta\bar{\theta}\bar{\theta}\square A(x) \\ &\quad + \sqrt{2}\theta\psi(x) - i/\sqrt{2}\theta\theta\partial_\mu \psi(x)\sigma^\mu \bar{\theta} + \theta\theta F(x) \end{aligned}$$

$$\theta^2 = \theta_1\theta_2, \quad \theta_1^2 = \theta_2^2 = 0!$$

component fields

# Gauge superfields

Gauge superfield

$$\begin{aligned}
 V(x, \theta, \bar{\theta}) = & C(x) + i\theta\chi(x) - i\bar{\theta}\bar{\chi}(x) + i\theta\theta M(x) - i\bar{\theta}\bar{\theta}M^+(x) \\
 & -\theta\sigma^\mu\bar{\theta}v_\mu(x) + i\theta\theta\bar{\theta}[\bar{\lambda}(x) + i\bar{\sigma}^\mu\partial_\mu\chi(x)] - i\bar{\theta}\bar{\theta}\theta[\lambda(x) + i\sigma^\mu\partial_\mu\bar{\chi}(x)] \\
 & + \frac{1}{2}\theta\theta\bar{\theta}\bar{\theta}[D(x) + \frac{1}{2}\square C(x)]
 \end{aligned}$$

Gauge transformation

$$\begin{aligned}
 C &\rightarrow C + A + A^* \\
 \chi &\rightarrow \chi - i\sqrt{2}\psi \\
 M &\rightarrow M - 2iF \\
 v_\mu &\rightarrow v_\mu - i\partial_\mu(A - A^*) \\
 \lambda &\rightarrow \lambda \\
 D &\rightarrow D
 \end{aligned}$$

$$V \rightarrow V + \Phi + \bar{\Phi}$$

Wess-Zumino gauge

$$C = \chi = M = 0$$

physical fields

Field strength tensor

$$W_\alpha = -\frac{1}{4}\bar{D}^2 e^V D_\alpha e^{-V}$$

Covariant derivatives

$$\begin{aligned}
 D_\alpha &= \frac{\partial}{\partial\theta^\alpha} + i\sigma_{\alpha\dot{\alpha}}^\mu \bar{\vartheta}^{\dot{\alpha}} \partial_\mu \\
 \bar{D}_{\dot{\alpha}} &= -\frac{\partial}{\partial\bar{\theta}^{\dot{\alpha}}} - i\theta_{\dot{\alpha}\alpha} \sigma_{\alpha\dot{\alpha}}^\mu \partial_\mu
 \end{aligned}$$

$$W_\alpha = -i\lambda_\alpha + \theta_\alpha D - \frac{i}{2}(\sigma^\mu \bar{\sigma}^\nu \theta)_\alpha F_{\mu\nu} + \theta^2 \sigma^\mu D_\mu \bar{\lambda}$$

# How to write SUSY Lagrangians

1<sup>st</sup> step

Take your favorite Lagrangian written in terms of fields

2<sup>nd</sup> step

**Replace** *Field*  $(\varphi, \psi, A_\mu) \Rightarrow$  *Superfield*  $(\Phi, V)$

3<sup>rd</sup> step

**Replace**

$$\textit{Action} = \int d^4x L(x) \quad \Rightarrow \quad \int d^4x d^4\theta L(x, \theta, \bar{\theta})$$

Grassmannian integration in superspace

$$\int d\theta_\alpha = 0, \quad \int \theta_\beta d\theta_\alpha = \delta_{\alpha\beta}$$



# Superfield Lagrangians

$$\text{Action} = \int d^4x L \quad \Rightarrow \quad \int d^4x d^4\theta L$$

Grassmannian integration in superspace

$$\int d\theta_\alpha = 0, \quad \int \theta_\beta d\theta_\alpha = \delta_{\alpha\beta}$$

Matter fields

$$L = \int d^2\theta d^2\bar{\theta} \Phi_i^+ \Phi_i + \int d^2\theta \underbrace{(\lambda_i \Phi_i + \frac{1}{2} m_{ij} \Phi_i \Phi_j + \frac{1}{3} y_{ijk} \Phi_i \Phi_j \Phi_k)}_{\text{Superpotential}} + h.c.]$$

Gauge fields

Superpotential

$$L = \frac{1}{4} \int d^2\theta W^\alpha W_\alpha + \int d^2\bar{\theta} \bar{W}^{\dot{\alpha}} \bar{W}_{\dot{\alpha}} = \frac{1}{2} D^2 - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - i\lambda \sigma^\mu D_\mu \bar{\lambda}$$

Gauge transformation  $\Phi \rightarrow e^{-ig\Lambda} \Phi, \Phi^+ \rightarrow \Phi^+ e^{ig\Lambda^+}, V \rightarrow V + i(\Lambda - \Lambda^+)$

Gauge invariant interaction

$$\Phi^+ \Phi \rightarrow \Phi^+ e^{gV} \Phi$$

# Gauge Invariant SUSY Lagrangian

Superfields

$$L_{SUSY\ YM} = \frac{1}{4} \int d^2\theta \text{Tr}(W^\alpha W_\alpha) + \frac{1}{4} \int d^2\bar{\theta} \text{Tr}(\bar{W}^\alpha \bar{W}_\alpha) \\ + \int d^2\theta d^2\bar{\theta} \bar{\Phi}_{ia} (e^{gV})^a_b \Phi_i^b + \int d^2\theta \mathcal{W}(\Phi_i) + \int d^2\bar{\theta} \bar{\mathcal{W}}(\bar{\Phi}_i)$$

Components

$$L_{SUSY\ YM} = -\frac{1}{4} F_{\mu\nu}^a F^{a\mu\nu} - i\lambda^a \sigma^\mu D_\mu \bar{\lambda}^a + \frac{1}{2} \underline{D^a D^a} \\ + (\partial_\mu A_i - igv_\mu^a T^a A_i)^\dagger (\partial_\mu A_i - igv_\mu^a T^a A_i) - i\bar{\psi}_i \sigma^\mu (\partial_\mu \psi_i - igv_\mu^a T^a \psi_i) \\ - \underline{D^a} g A_i^\dagger T^a A_i - i\sqrt{2} g A_i^\dagger T^a \lambda^a \psi_i + i\sqrt{2} g \bar{\psi}_i T^a \bar{\lambda}^a A_i + \underline{F_i^\dagger F_i} \\ + \frac{\partial \mathcal{W}}{\partial A_i} \underline{F_i} + \frac{\partial \bar{\mathcal{W}}}{\partial A_i^\dagger} \underline{F_i^\dagger} - \frac{1}{2} \frac{\partial^2 \mathcal{W}}{\partial A_i \partial A_j} \psi_i \psi_j - \frac{1}{2} \frac{\partial^2 \bar{\mathcal{W}}}{\partial A_i^\dagger \partial A_j^\dagger} \bar{\psi}_i \bar{\psi}_j$$

Potential

$$D^a = -g A_i^\dagger T^a A_i, \quad F_i = -\frac{\partial \mathcal{W}}{\partial A_i} \quad \rightarrow \quad V = \frac{1}{2} D^a D^a + F_i^\dagger F_i$$