

The linear dilaton: from the clockwork mechanism to its supergravity embedding

Chrysoula Markou

August 30, 2019
SQS '19, Yerevan

Max-Planck-Institut
für Physik



Overview

1 The clockwork mechanism

- The clockwork scalar
- Motivation and basic references
- The clockwork graviton
- Continuum

2 Little String Theory

- Definition
- Geometry
- Phenomenology

3 Supergravity embedding

- Structure of $\mathcal{N} = 2, D = 5$ Supergravity
- Effective supergravity

Setup

- ➊ massless scalar $\pi(x)$, symmetry $U(1)$
- ➋ $N+1$ copies: $\pi_j(x)$, $j = 0, 1, \dots, N$, symmetry $U(1)^{N+1}$ (at least)
- ➌ explicit breaking $U(1)^{N+1} \rightarrow U(1)$ via a mass-mixing that involves **near-neighbours** only

$$\mathcal{L} = -\frac{1}{2} \sum_{j=0}^N \partial_\mu \pi_j \partial^\mu \pi_j - \frac{1}{2} m^2 \sum_{j=0}^{N-1} (\pi_j - \textcolor{red}{q} \pi_{j+1})^2 + \dots$$

q treats the site j and the site $j+1$ asymmetrically



Eigenmodes $\pi = Oa$

① massless Goldstone a_0 , $O_{j0} = \frac{N_0}{q^j}$

② massive Clockwork gears a_k , $O_{jk} = \mathcal{N}_k \left[q \sin \frac{jk\pi}{N+1} - \sin \frac{(j+1)k\pi}{N+1} \right]$

now couple eg. topological term to the **N -th site only**

$$\frac{1}{f} \pi_N G_{\mu\nu} \tilde{G}^{\mu\nu} = \left\{ \frac{1}{f_0} a_0 - \sum_{k=1}^N (-)^k \frac{1}{f_k} a_k \right\} G_{\mu\nu} \tilde{G}^{\mu\nu}$$

① $f_0/f \sim q^N$: exponential enhancement

② $f_k/f \sim N^{3/2}/k$: mild N dependance

- the clockwork as a *renormalizable* scalar field theory:
[Choi and Im '15](#)
[Kaplan and Rattazzi '15](#)
(see also natural inflation and relaxion models)
- the clockwork as an *effective* low-energy field theory:
[Giudice and McCullough '16](#)
 - Recipe: $N + 1$ of particle P whose masslessness is protected by symmetry S & explicit breaking via near-neighbour interactions
 - Known examples for the following spins: 0, 1/2, 1, 2
- growing pheno literature (LHC signatures, inflation, ...)
- group theory of the clockwork:
[Craig, Garcia Garcia, Sutherland '17](#)
[Giudice and McCullough '17](#)

- ① @ the linear level: use Fierz–Pauli 1939 mass term

$$\mathcal{L}_{int} = -\frac{1}{2}m^2 \sum_{j=0}^{N-1} \left[(h_j^{\mu\nu} - \textcolor{red}{q} h_{j+1}^{\mu\nu})^2 - [\eta_{\mu\nu}(h_j^{\mu\nu} - \textcolor{red}{q} h_{j+1}^{\mu\nu})]^2 \right]$$

Giudice and McCullough '16

- ② @ the nonlinear level: use Hinterbichler and Rosen '12 interaction

$$\mathcal{S}_V = - \sum_{i,j,k,l} \int \textcolor{red}{T}_{ijkl} \varepsilon_{abcd} (E_i)^a \wedge (E_j)^b \wedge (E_k)^c \wedge (E_l)^d$$

Niedermann, Padilla, Saffin '18

in both cases: *asymmetric* gear shifts $\xi_i^\mu = \frac{1}{q^i} \xi^\mu$

Setup

discrete and finite lattice dimension \rightarrow continuous y on S^1/\mathbb{Z}_2

Giudice and McCullough '16

- ① Consider real scalar ϕ in 5D with metric Ansatz:

$$ds^2 = X(|y|)dx^2 + Y(|y|)dy^2$$

- ② Discretize: $y \rightarrow y_j = ja$, $\phi, X, Y \rightarrow \phi_j, X_j, Y_j$

$$\mathcal{L} = -\frac{1}{2} \sum_{j=0}^N \partial_\mu \phi_j \partial^\mu \phi_j - \frac{1}{2} \sum_{j=0}^{N-1} \frac{X_j}{a^2 Y_j} \left(\phi_j - \frac{X_j^{1/2} Y_j^{1/4}}{X_{j+1}^{1/2} Y_{j+1}^{1/4}} \phi_{j+1} \right)^2$$

- ③ Identify: $X_j \sim Y_j \sim e^{-\frac{4}{3} \frac{k\pi R}{N} j}$ with $q = e^{\textcolor{blue}{k}a}$

Eigenmodes

① zero-mode $\psi_0 \sim e^{-k|y|}$

② Kaluza–Klein excitations $\psi_n \sim \left(\frac{kR}{n} \sin \frac{n|y|}{R} + \cos \frac{n|y|}{R} \right)$

now couple eg. topological term to $y = 0$ only
 (flipped k sign for pheno reasons)

$$\frac{1}{f} \int dy \delta(y) \phi G_{\mu\nu} \tilde{G}^{\mu\nu} = \left\{ \frac{1}{f_0} \hat{\phi}_0 + \sum_{n=1}^N \frac{1}{f_n} \hat{\phi}_n \right\} G_{\mu\nu} \tilde{G}^{\mu\nu}$$

① $f_0/f \sim e^{k\pi R}/\sqrt{k\pi R}$: exponential enhancement

② $f_n \sim f$: roughly equal

Origin and comparison

- Continuum clockwork
 - 5D metric: $ds^2 = e^{\frac{4}{3}k|y|}(dx^2 + dy^2)$
 - KK spectrum: $m_n^2 = k^2 + \left(\frac{n}{R}\right)^2$
- ⇒ precisely those of the 5D toy model for the holographic dual of 6D Little String Theory as we will see now!
- Discretize 5th dimension of other models:
 - large extra dimensions: no warping and *no* “clockworking”
 - Randall–Sundrum: warping and “clockworking”
 - but \nexists simple one-parameter deformation from the clockwork to Randall–Sundrum
- to explain eg. hierarchy, need only $\mathcal{O}(10)$ sites and $q \sim 4$

- there exist two inequivalent limits of string theory:

① $E \ll m_S$: low-energy limit

② $g_S \rightarrow 0$: “little string theories”

Berkooz, Rozali, Seiberg '97, Seiberg '97, ...

- simple example of LST: take $g_S \rightarrow 0$ for a stack of N $NS5$ -branes in type IIB

① m_S and N are the only parameters

② bulk dynamics decouple

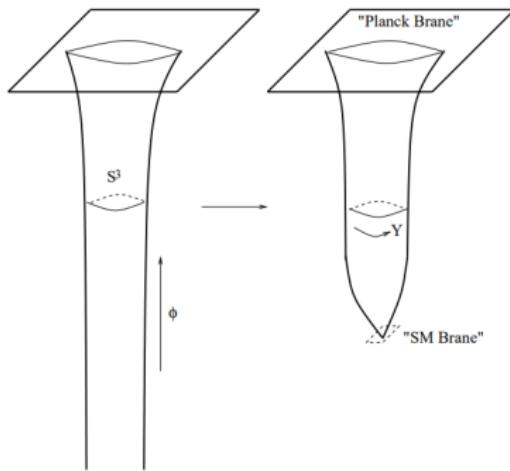
- Properties of LST

① non-gravitational, since $M_{Pl} \rightarrow \infty$ as $g_S \rightarrow 0$

② non-local, eg. Hagedorn density of states at high energies, T-duality

③ interacting, with $\frac{1}{g_N^2} = m_S^2$

- “holographic” dual of 6D LST: string theory on $\mathbb{R}^{5,1} \times \mathbb{R}_y \times S_N^3$
Aharony, Berkooz, Kutasov, Seiberg ’98
 - infinite throat connected to asymptotically flat spacetime
 - linear dilaton $\Phi = -\frac{m_S}{\sqrt{N}}y$
- compactify y on interval (therby treating strong coupling problem and obtaining a large but finite M_{Pl} in the weak coupling regime): cigar-like geometry, Antoniadis, Dimopoulos, Giveon ’01



- 5D graviton–dilaton toy model for cigar: Antoniadis, Arvanitaki, Dimopoulos, Giveon '11

$$e^{-1} \mathcal{L}_{LST} = e^{-\frac{\Phi}{M_5^{3/2}}} \left(M_5^3 \mathcal{R} + (\nabla \Phi)^2 - \Lambda \right) + \text{branes}$$

- impose linear dilaton background $\Phi = \alpha|y|$:
 - background metric: $ds^2 = e^{-\frac{2}{3}\alpha|y|}(dx^2 + dy^2)$
 - KK spectrum: $m_n^2 = \left(\frac{\alpha}{2}\right)^2 + \left(\frac{n\pi}{r_c}\right)^2$
 \Rightarrow same as continuum clockwork!
- supergravity embedding
 Kehagias, Riotto '17, Antoniadis, Delgado, CM, Pokorski '17

$\mathcal{N} = 2$, $D = 5$ Supergravity

- pure supergravity multiplet: (e_M^m, ψ_M^i, A_M^0)
- vector multiplet: (ϕ, λ^i, A_M^1)
- Maxwell–Einstein supergravity: $(e_M^m, \psi_M^i, A_M^I, \lambda_i^a, \phi^x)$
 - ① $\{\phi^x\}$: coordinates of \mathcal{M}
 - ② $\{\lambda_i^a\}$: tangent space group $SO(n)$
- Lagrangian:

$$\begin{aligned}
 e^{-1}\mathcal{L}_{ME} &= \frac{1}{2\kappa^2} \left[R(\omega) - \overline{\psi}_M^i \Gamma^{MNP} D_N \psi_{Pi} \right] - \frac{1}{4} G_{IJ} F_{MN}^I F^{MNJ} \\
 &\quad - \frac{1}{2} \overline{\lambda}^{ia} (\not{D} \delta^{ab} + \Omega_x{}^{ab} \not{\partial} \phi^x) \lambda_i^b - \frac{1}{2} g_{xy} (\partial_M \phi^x) (\partial^M \phi^y) \\
 &\quad - \frac{i}{2} \overline{\lambda}^{ia} \Gamma^M \Gamma^N \psi_{Mi} f_x^a \partial_N \phi^x + \frac{1}{4} \textcolor{red}{h}_I^a \overline{\lambda}^{ia} \Gamma^M \Gamma^{\Lambda P} \psi_{Mi} F_{\Lambda P}^I \\
 &\quad + \frac{i\kappa}{4} \Phi_{Iab} \overline{\lambda}^{ia} \Gamma^{MN} \lambda_i^b F_{MN}^I + \frac{\kappa e^{-1}}{6\sqrt{6}} C_{IJK} \epsilon^{MNP\Sigma\Lambda} F_{MN}^I F_{P\Sigma}^J A_\Lambda^K \\
 &\quad - \frac{3i}{8\kappa\sqrt{6}} \textcolor{red}{h}_I \left(\overline{\psi}_M^i \Gamma^{MNP\Sigma} \psi_{Ni} F_{P\Sigma} + 2 \overline{\psi}^{Mi} \psi_i^N F_{MN} \right) + \dots
 \end{aligned}$$

Günaydin–Sierra–Townsend '84

Constraints

① Algebraic:

$$h_x^I h_I = h_{Ix} h^I = 0 \quad , \quad h_x^I h_y^J G_{IJ} = g_{xy}$$

$$G_{IJ} = h_I h_J + h_I^x h_J^y g_{yx} = h_I h_J + h_I^x h_{Jx}$$

$$h^I h_I = 1 \quad , \quad h_x^I h_I^y = \delta_x^y$$

$$\Rightarrow \quad h^I G_{IJ} = h_J \quad , \quad h_x^I G_{IJ} = h_{Jx}$$

② Differential:

$$h_{I,x} = \sqrt{\frac{2}{3}} h_{Ix} \quad , \quad h_{,x}^I = -\sqrt{\frac{2}{3}} h_x^I$$

$$h_{Ix;y} = \sqrt{\frac{2}{3}} (g_{xy} h_I + T_{xyz} h_I^z) \quad , \quad h_{x;y}^I = -\sqrt{\frac{2}{3}} (g_{xy} h^I + T_{xyz} h^{Iz})$$

Embedding

- \mathcal{E} with coordinates $X^I = X^I(\phi^x, \mathcal{N})$
- hypersurfaces \mathcal{M}_k defined by $\ln \mathcal{N} = k$ with normal vectors
 $n_I = \partial_I \ln \mathcal{N}$
- **orthonormality** relations:

$$X_{,x}^I n_I = 0 \quad , \quad \frac{3}{2} \beta^2 X_{,x}^I X_I^{,y} = \delta_x^y$$

$$\Rightarrow \text{ identify: } h_I \stackrel{!}{=} \alpha n_I \quad , \quad h^I \stackrel{!}{=} \beta X^I$$

- can also show that:

$$\mathcal{N} = \beta^3 C_{IJK} X^I X^J X^K$$

- \mathcal{M} : special case $k = 0 \Rightarrow \beta^3 C_{IJK} X^I X^J X^K = 1$

Gauging

- gauge $U(1)$ subgroup of $SU(2)$ w.r.t. $A_M = v_I A_M^I$
- new Lagrangian terms:

$$\begin{aligned} e^{-1}\mathcal{L}' &= -\frac{g^2}{\kappa^4} \textcolor{red}{P} - \frac{i\sqrt{6}}{8} \frac{g}{\kappa^3} \overline{\psi}_M^i \Gamma^{MN} \psi_N^j \delta_{ij} P_0 \\ &\quad - \frac{1}{\sqrt{2}} \frac{g}{\kappa^2} \overline{\lambda}^{ia} \Gamma^M \psi_M^j \delta_{ij} P_a + \frac{i}{2\sqrt{6}} \frac{g}{\kappa} \overline{\lambda}^{ia} \lambda^{jb} \delta_{ij} P_{ab} \end{aligned}$$

- new terms in the fermion transformations:

$$\begin{aligned} \delta' \psi_{Mi} &= \frac{i}{2\kappa\sqrt{6}} g P_0 \Gamma_M \epsilon_{ji} \delta^{jk} \epsilon_k \\ \delta' \lambda_i^a &= \frac{1}{\kappa^2 \sqrt{2}} g P^a \epsilon_{ji} \delta^{jk} \epsilon_k \end{aligned}$$

- algebraic and differential constraints

Günaydin–Sierra–Townsend '84, '85

Embedding

- use a $U(1)$ gauging of $\mathcal{N} = 2$, $D = 5$ Maxwell–Einstein supergravity with *one* vector multiplet

- Ansatz:**

$$\mathcal{N} = st^2 + at^3 \quad , \quad X^0 \equiv s \, , \, X^1 = t$$

Klemm, Lerche, Mayr '95, Antoniadis, Ferrara, Taylor, '95

- calculate potential:

$$P = -3A_P \left(\frac{A_P}{4}t^2 + B_P \frac{1}{t} \right)$$

- redefine $\sqrt{3} \ln t = \Phi$, set $\frac{3}{4}g^2 A_P^2 = -\Lambda$, $B_P = 0$

$$\Rightarrow e^{-1}\mathcal{L}_{LST} = \frac{1}{2}\mathcal{R} - \frac{1}{2}(\partial_M \Phi)(\partial^M \Phi) - e^{\frac{2}{\sqrt{3}}\Phi} \Lambda$$

Partial breaking

- $\Phi = Cy$, fine-tuning condition: $C = \pm \frac{gA_P}{\sqrt{2}}$
- fermion transformations:

$$\tilde{\delta}\psi_{\mu i} = \frac{i}{2\sqrt{3}}\Gamma_\mu \left(iC\Gamma^5\epsilon_i + \frac{gA_P}{\sqrt{2}}\varepsilon_{ji}\delta^{jk}\epsilon_k \right)$$

$$\tilde{\delta}\lambda_i = -\frac{1}{2}e^{\frac{1}{\sqrt{3}}Cy} \left(iC\Gamma^5\epsilon_i + \frac{gA_P}{\sqrt{2}}\varepsilon_{ji}\delta^{jk}\epsilon_k \right)$$

so that

$$\tilde{\delta}(\lambda_1 - i\Gamma^5\lambda_2) = 0$$

$$\tilde{\delta}(\lambda_1 + i\Gamma^5\lambda_2) \sim \epsilon_2 - i\Gamma^5\epsilon_1$$

\Rightarrow partial supersymmetry breaking

- three free parameters: g, v_0, v_1 . Mass gap: $\sim g$

Summary

- Gauged $\mathcal{N} = 2$, $D = 5$ supergravity can accommodate the linear dilaton model (**minimal case**)
- Partial breaking *due to the linear dilaton bkg*
- 5D supergravity embedding of other models
Im, Nilles, Olechowski '18
- Branes **compatible** with the unbroken supersymmetry
Antoniadis, Delgado, CM, Pokorski (ongoing work)
- deeper link between LST and the clockwork?