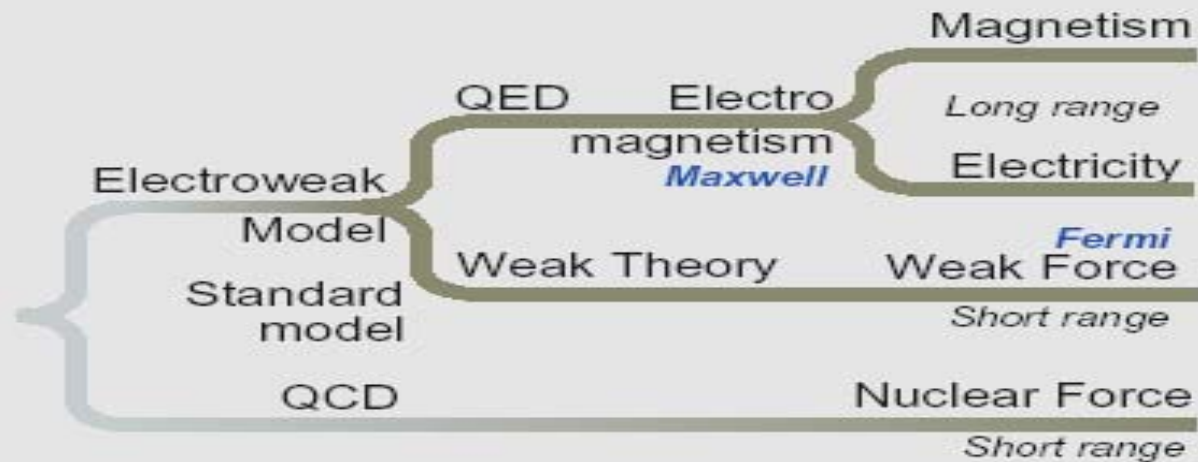
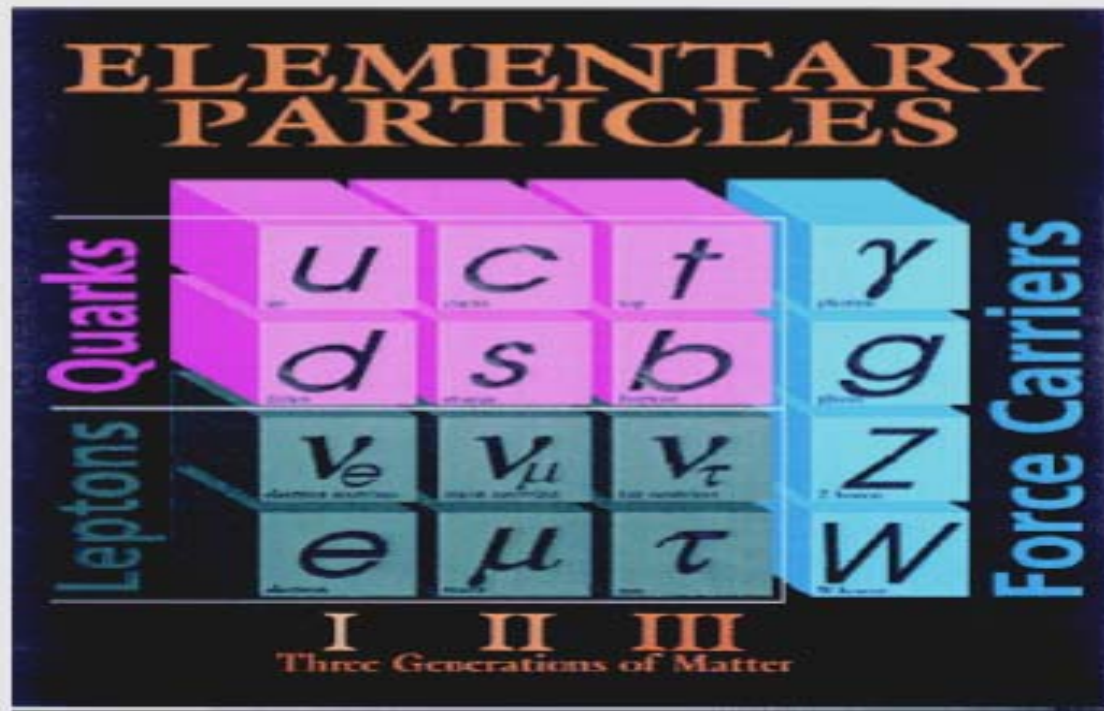


**ROUND TABLE ITALY-RUSSIA,  
DUBNA, DECEMBER 19, 2009**

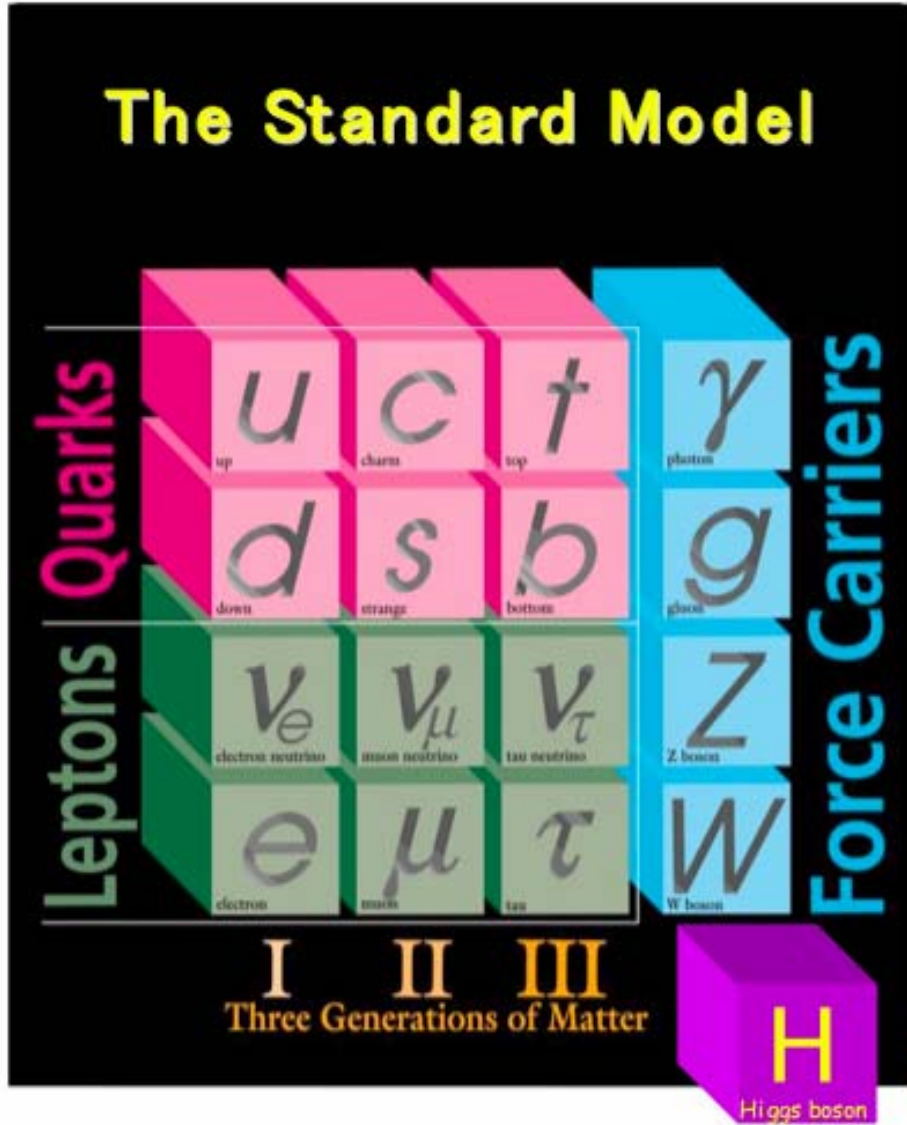
**On the DARK MATTER  
at the time of LHC**

**Antonio Masiero  
Univ. of Padova  
and  
INFN, Padova**

# THE G-W-S STANDARD MODEL

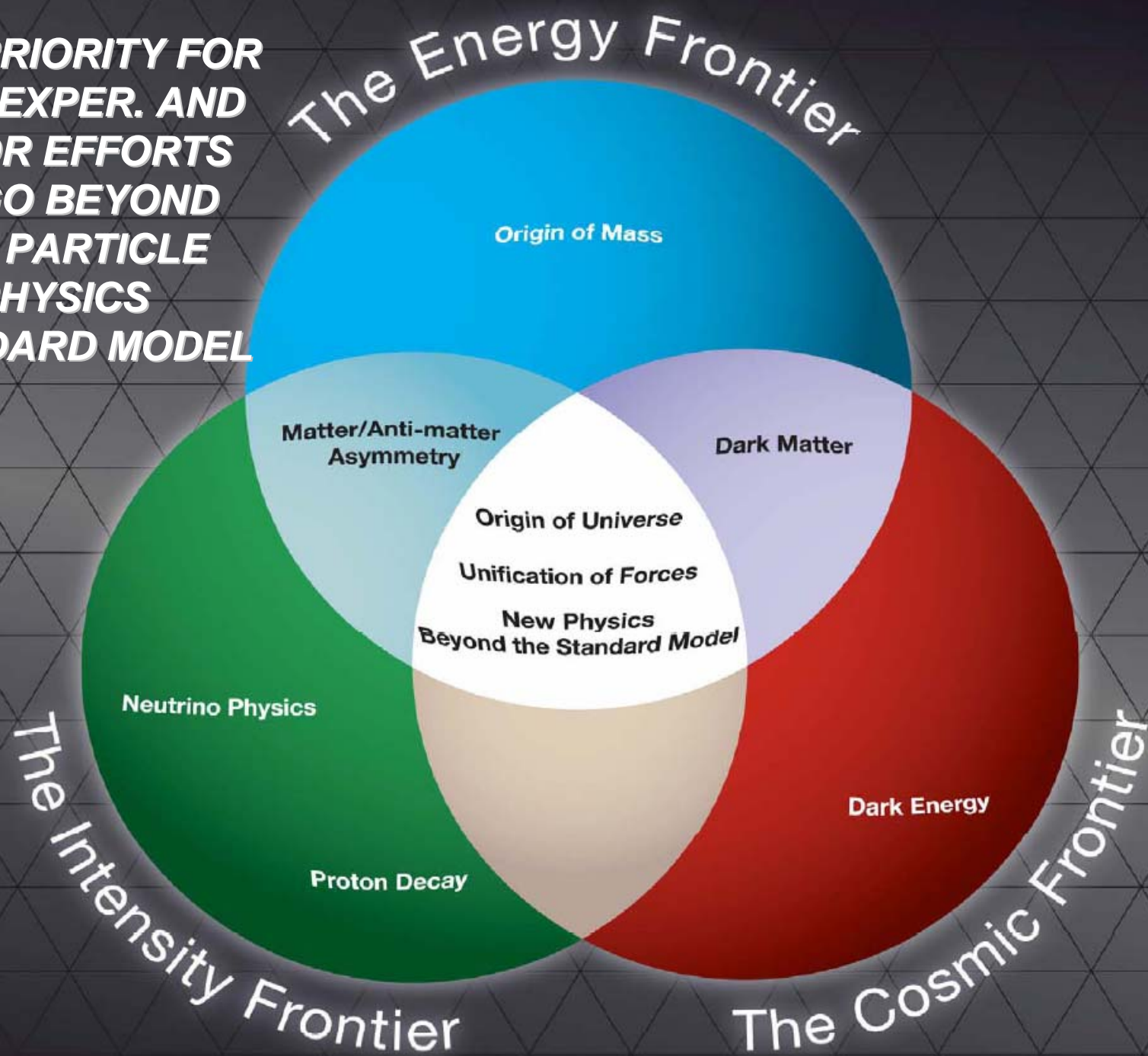


Gravity  
?



?

**HIGH PRIORITY FOR  
INFN: EXPER. AND  
THEOR EFFORTS  
TO GO BEYOND  
THE PARTICLE  
PHYSICS  
STANDARD MODEL**





## The muon $g-2$ : Standard Model vs. Experiment

- Adding up all the above contribution we get the following SM predictions for  $a_\mu$  and comparisons with the measured value:

	$a_\mu^{\text{SM}} \times 10^{11}$	$\Delta a_\mu \times 10^{11}$	$\sigma$
[1]	116 591 793 (60)	287 (87)	3.3
[2]	116 591 778 (61)	302 (88)	3.4
[3]	116 591 807 (72)	273 (96)	2.8
[4]	116 591 828 (63)	252 (89)	2.8
[5]	116 591 991 (70)	89 (95)	0.9

with  $a_\mu^{\text{H}^0(|b|)} = 110 (40) \times 10^{-11}$ .

$\Delta a_\mu = a_\mu^{\text{EXP}} - a_\mu^{\text{SM}}$ .



2

- [1] Eidelman at ICHEP06 & Davier at TAU06 (update of ref. [5]).
- [2] Hagiwara, Martin, Nomura, Teubner, PLB649 (2007) 173.
- [3] F. Jegerlehner, PhiPsi 08, Frascati, April 2008.
- [4] J.F. de Troconiz and F.J. Yndurain, PRD71 (2005) 073008.
- [5] Davier, Eidelman, Hoecker and Zhang, EPJC31 (2003) 503 ( $\tau$  data).

- The th error is now the same (or even smaller) as the exp. one!
- If BaBar's prelim. results are used instead,  $\Delta a_\mu$  drops to  $\sim 1.7\sigma$ .

Courtesy of M. Passera

**WG on Rad. Corrections and MC Generators for Low Energies**

( coll. with Novosibirsk and Dubna)

Coll. Agreement INFN - Moscow State Univ. for theor, research in this field

# Present “Observational” Evidence for New Physics

- **NEUTRINO MASSES** 
- **DARK MATTER** 
- **MATTER-ANTIMATTER ASYMMETRY** 
- **INFLATION** 

# **EVIDENCE FOR DM**

**Various astrophysical sources have confirmed the existence of Dark Matter (DM)**

- **Binding of Galaxies in Clusters ( F. Zwicky, 1933 )**
- **Rotation curves of Galaxies ( V.C. Rubin and W.K. Ford, 1970 )**
- **Bindings of hot gases in clusters**
- **Gravitational Lensing observations**
- **Large Scale Structure simulations**
- **High z - Supernovae**
- **Observations of colliding clusters of Galaxies**

**The most direct and accurate evidence comes from WMAP by measuring anisotropies of the CMB power spectrum**

**~ 73% DarkEnergy, ~ 23% DarkMatter, 4% Baryons**

**DM → NEW PHYSICS BEYOND THE**  
**( PARTICLE PHYSICS ) SM - if Newton is right**  
**at scales > size of the Solar System**

- $\Omega_{\text{DM}} = 0.233 \pm 0.013$  \*

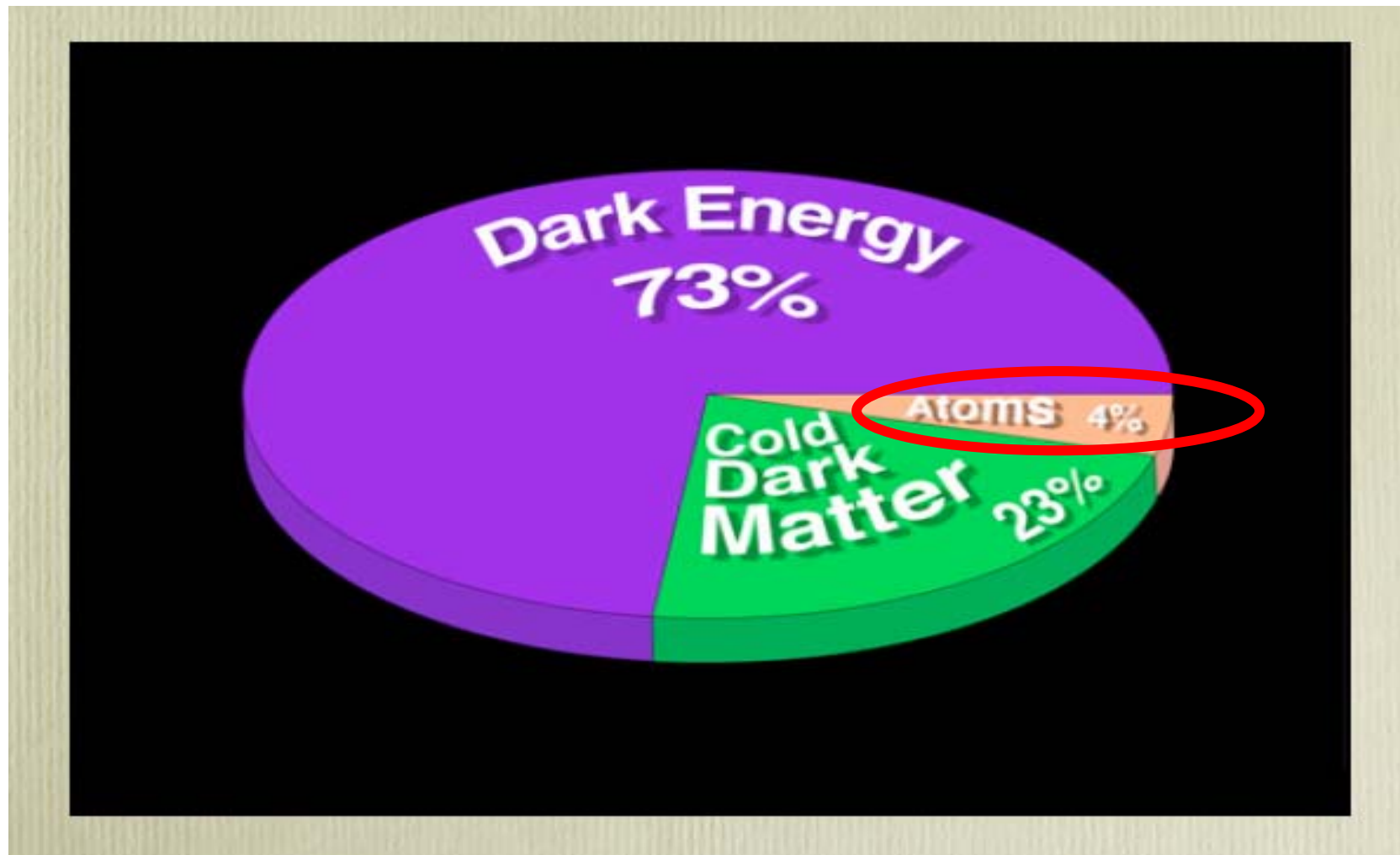
- $\Omega_{\text{baryons}} = 0.0462 \pm 0.0015$  \*\*

\*from CMB ( 5 yrs. of WMAP) + Type I  
Supernovae + Baryon Acoustic  
Oscillations (BAO)

\*\*CMB + Type I SN + BAO in agreement with  
Nucleosynthesis (BBN)



***In terms of energy content, the astonishing progress of human knowledge of the Universe has reached only 4% of its content!!***



# **STABLE ELW. SCALE WIMPs** ( **Weakly Interactive Massive Particle** )

<p>1) <b>ENLARGEMENT OF THE SM</b></p>	<p><b>SUSY</b> (<math>x^\mu, \theta</math>)</p> <p>Anticomm. Coord.</p>	<p><b>EXTRA DIM.</b> (<math>x^\mu, j^i</math>)</p> <p>New bosonic Coord.</p>	<p><b>LITTLE HIGGS.</b> SM part + new part</p> <p>to cancel <math>\Lambda^2</math> at 1-Loop</p>
<p>2) <b>SELECTION RULE</b></p>	<p><b>R-PARITY LSP</b></p>	<p><b>KK-PARITY LKP</b></p>	<p><b>T-PARITY LTP</b></p>
<p>→ <b>DISCRETE SYMM.</b></p>	<p>Neutralino spin 1/2</p>	<p>spin1</p>	<p>spin0</p>
<p>→ <b>STABLE NEW PART.</b></p>	<p>↓</p>	<p>↓</p>	<p>↓</p>
<p>3) <b>FIND REGION (S) PARAM. SPACE WHERE THE "L" NEW PART. IS NEUTRAL + <math>\Omega_L h^2</math> OK</b></p>	<p><math>m_{LSP}</math></p> <p>~100 - 200 GeV *</p>	<p><math>m_{LKP}</math></p> <p>~600 - 800 GeV</p>	<p><math>m_{LTP}</math></p> <p>~400 - 800 GeV</p>

\* But abandoning gaugino-masss unif. → Possible to have  $m_{LSP}$  down to 7 GeV

Bottino, Donato, Fornengo, Scopel

# HUMAN PRODUCTION OF WIMPs

## WIMPS HYPOTHESIS

DM made of particles with  
mass 10Gev - 1Tev

*ELW scale*

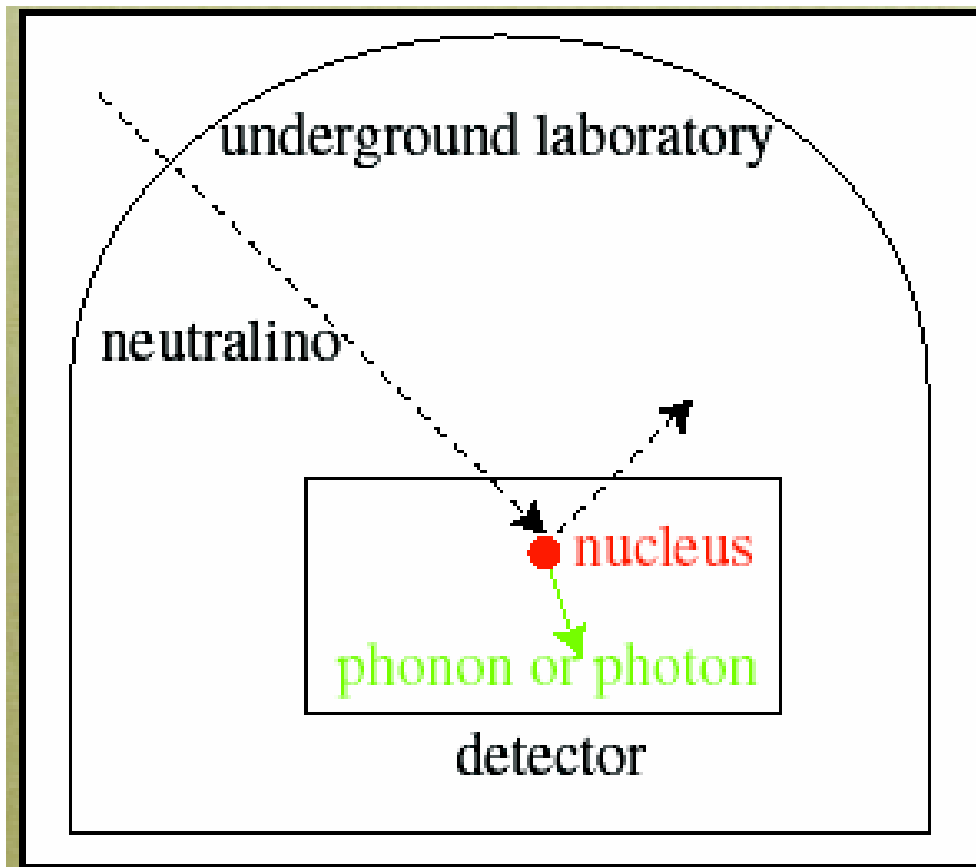
With **WEAK INTERACT.**

*LHC, ILC may  
PRODUCE WIMPS*

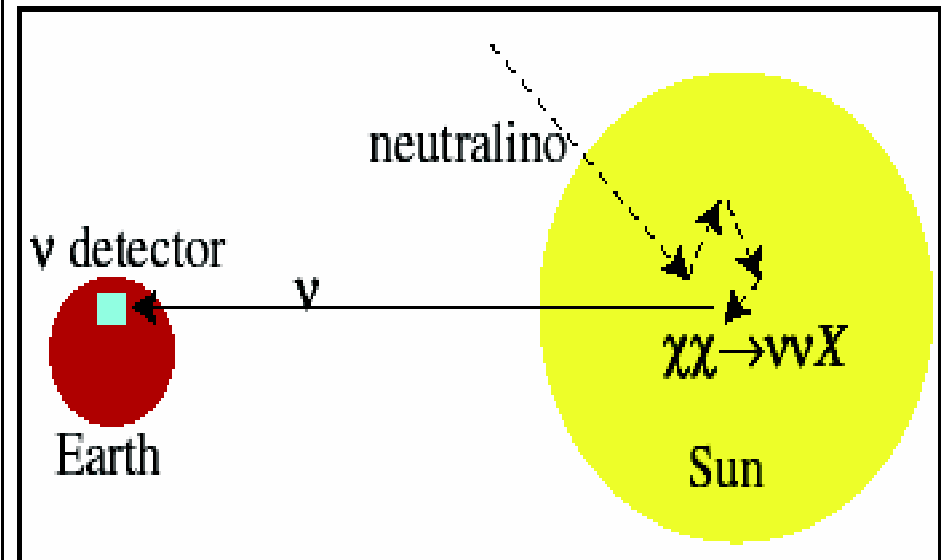
WIMPS escape the detector  
→ MISSING ENERGY  
SIGNATURE

***POSSIBILITY TO CREATE OURSELVES IN OUR  
ACCELERATORS THOSE DM PARTICLES WHICH  
ARE PART OF THE RELICS OF THE PRIMORDIAL  
PLASMA AND CONSTITUTE 1/4 OF THE WHOLE  
ENERGY IN THE UNIVERSE***

# HUNTING FOR DARK MATTER



**DIRECT DM SEARCHES**



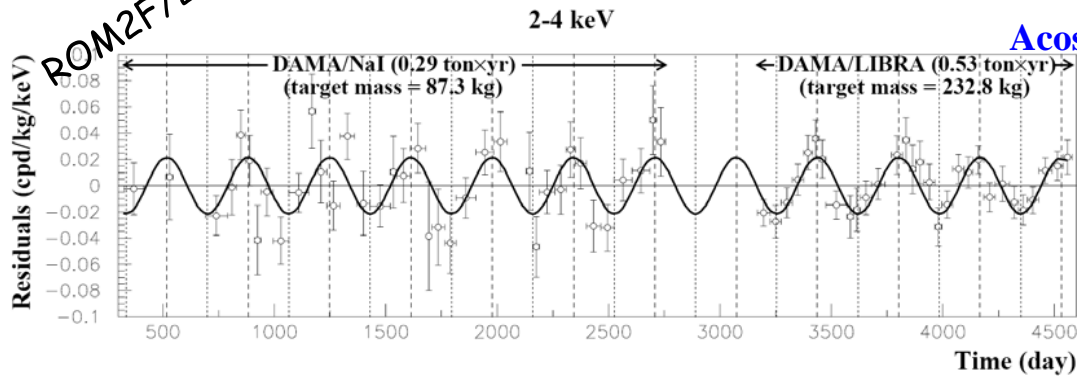
**INDIRECT DM SEARCHES**

# Model Independent Annual Modulation Result

DAMA/NaI (7 years) + DAMA/LIBRA (4 years) Total exposure: 300555 kg×day = 0.82 ton×yr  
 experimental single-hit residuals rate vs time and energy

ROM2F/2008/07

$\text{Acos}[\omega(t-t_0)]$ ; continuous lines:  $t_0 = 152.5$  d,  $T = 1.00$  y

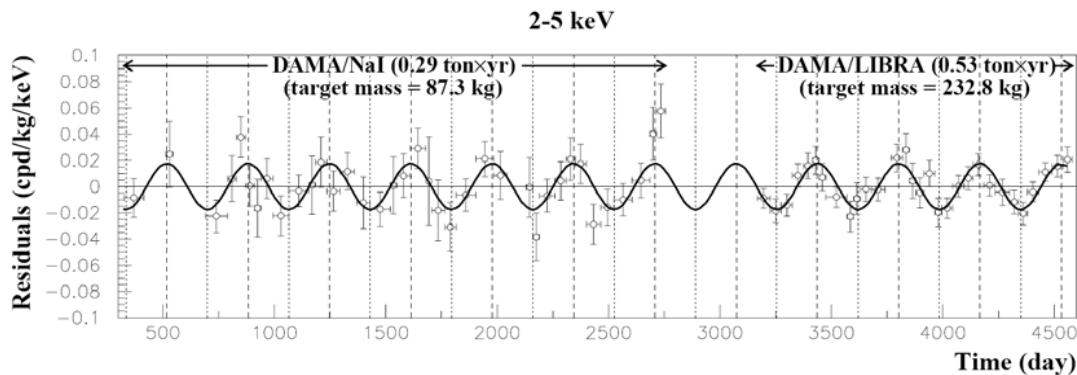


## 2-4 keV

$A = (0.0215 \pm 0.0026)$  cpd/kg/keV  
 $\chi^2/\text{dof} = 51.9/66$  **8.3  $\sigma$  C.L.**

Absence of modulation? No

$\chi^2/\text{dof} = 117.7/67 \Rightarrow P(A=0) = 1.3 \times 10^{-4}$

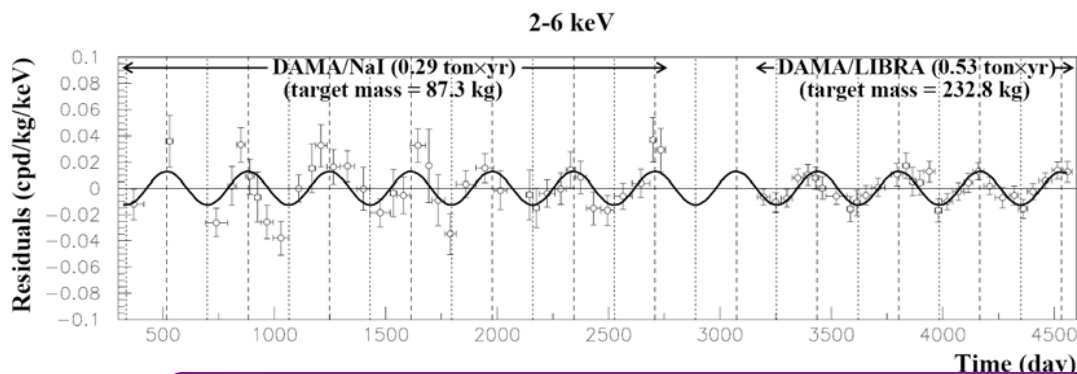


## 2-5 keV

$A = (0.0176 \pm 0.0020)$  cpd/kg/keV  
 $\chi^2/\text{dof} = 39.6/66$  **8.8  $\sigma$  C.L.**

Absence of modulation? No

$\chi^2/\text{dof} = 116.1/67 \Rightarrow P(A=0) = 1.9 \times 10^{-4}$



## 2-6 keV

$A = (0.0129 \pm 0.0016)$  cpd/kg/keV  
 $\chi^2/\text{dof} = 54.3/66$  **8.2  $\sigma$  C.L.**

Absence of modulation? No

$\chi^2/\text{dof} = 116.4/67 \Rightarrow P(A=0) = 1.8 \times 10^{-4}$

The data favor the presence of a modulated behavior with proper features at 8.2 $\sigma$  C.L.

# **The first glimpse of dark matter?**

By Victoria Gill

Science reporter, BBC News

Story from BBC NEWS:

<http://news.bbc.co.uk/go/pr/fr/-/2/hi/science/nature/8420089.stm>

**2009/12/18** 11:52:13 GMT

arXiv astro-ph 0912.3592v1  
CDMS EXP.

**We report results from a blind analysis of the nal data taken  
with the Cryogenic Dark Matter**

**Search experiment (CDMS II) at the Soudan Underground  
Laboratory, Minnesota, USA. A total**

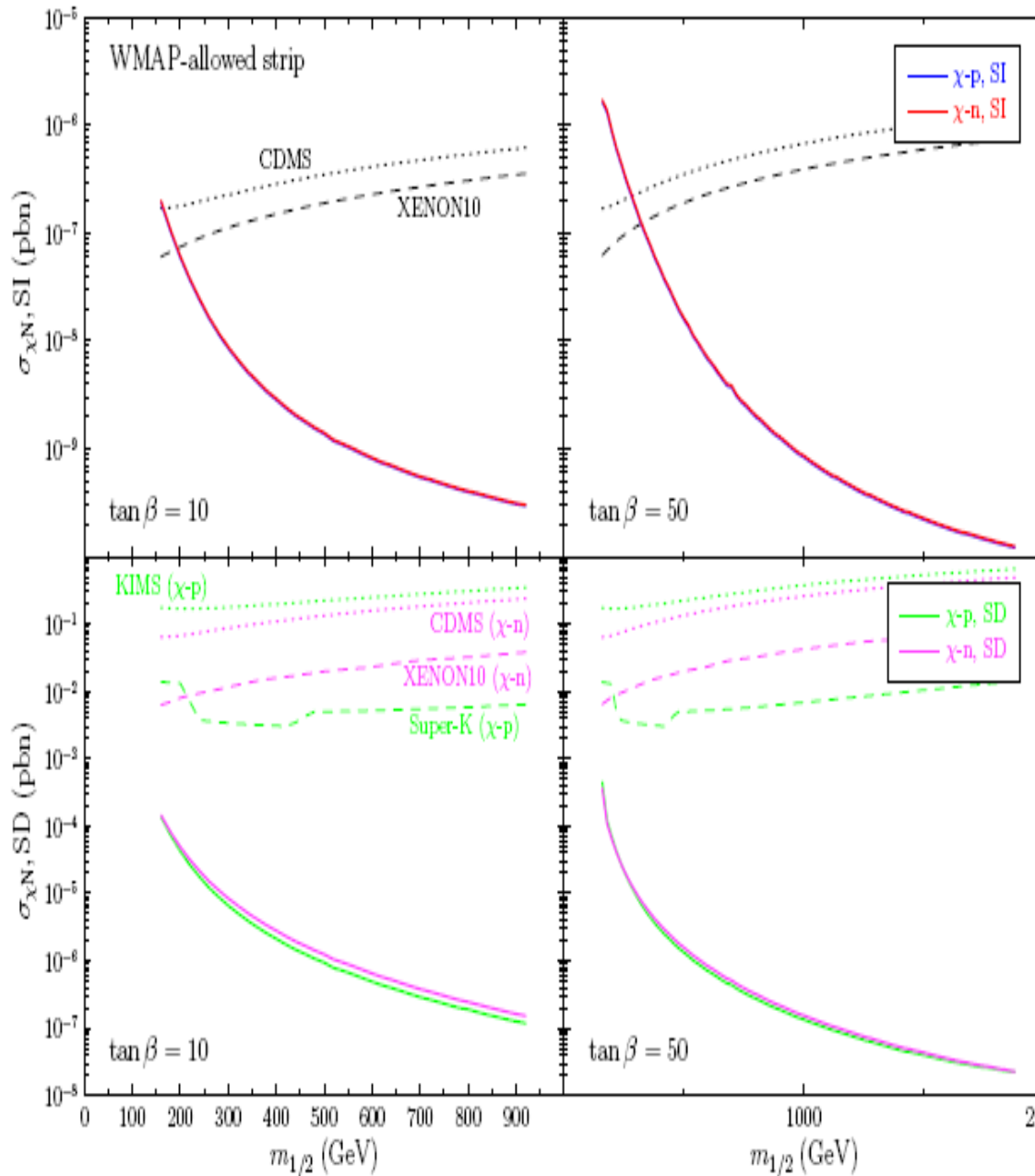
**raw exposure of 612 kg-days was analyzed for this work.**

**We observed two events in the signal region;  
based on our background estimate, the probability of  
observing two or more background events is 23%.**



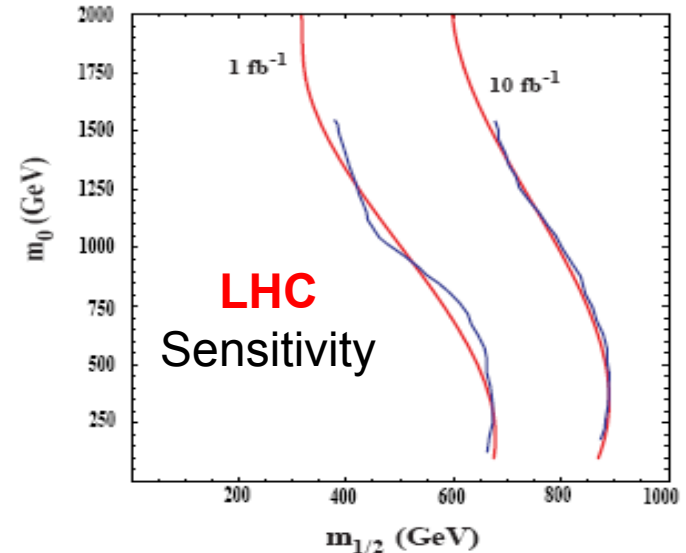
**Neutralino-nucleon scattering cross sections** along the WMAP-allowed coannihilation strip for  $\tan\beta=10$  and **coannihilation/funnel strip** for  $\tan\beta=50$  using the hadronic parameters

ELLIS. OLIVE. SAVAGE 



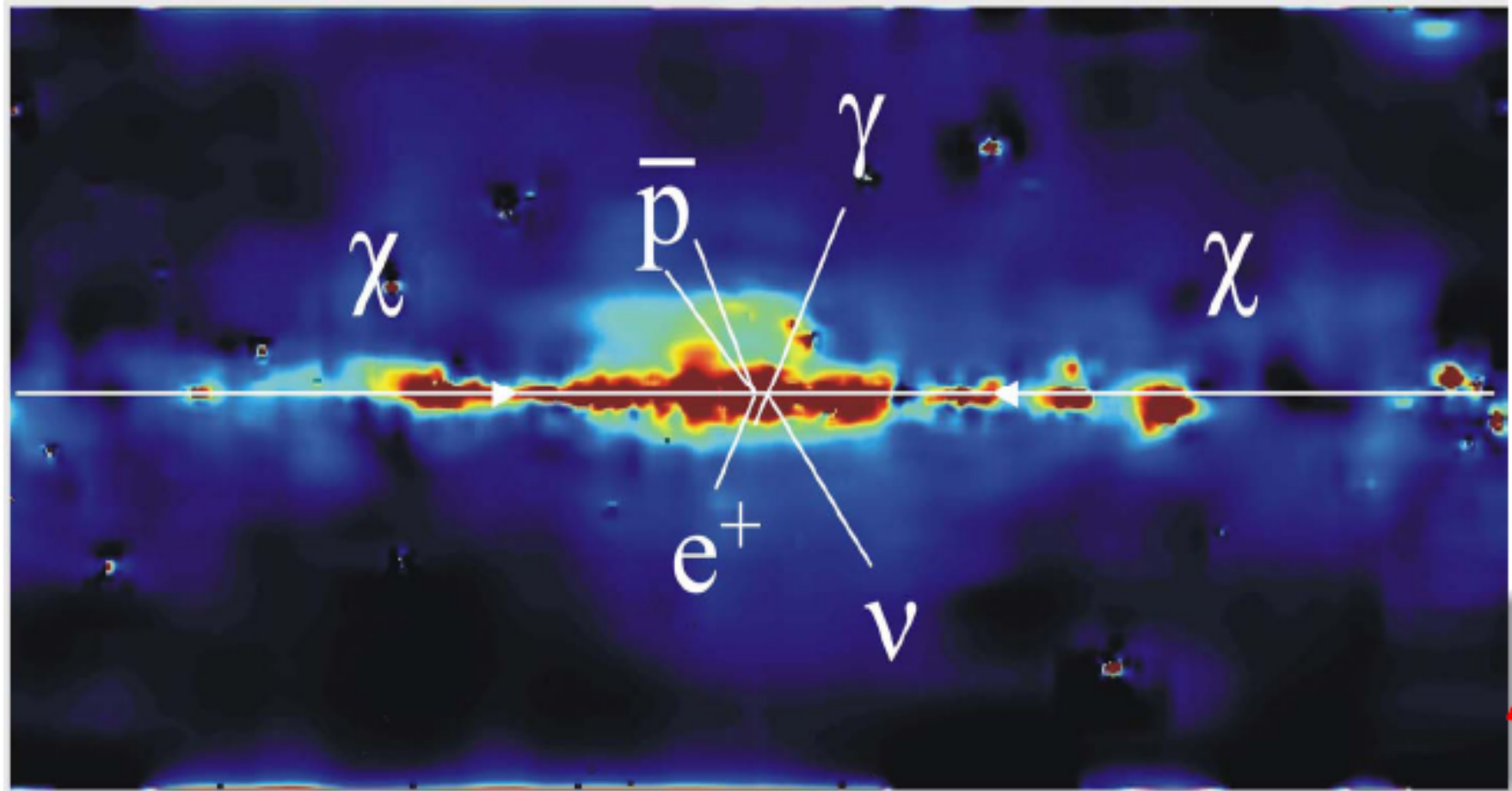
$m_u/m_d$	$0.553 \pm 0.043$
$m_d$	$5 \pm 2$ MeV
$m_s/m_d$	$18.9 \pm 0.8$
$m_c$	$1.25 \pm 0.09$ GeV
$m_b$	$4.20 \pm 0.07$ GeV
$m_t$	$171.4 \pm 2.1$ GeV
$\sigma_0$	$36 \pm 7$ MeV
$\Sigma_{\pi N}$	$64 \pm 8$ MeV
$a_3^{(p)}$	$1.2695 \pm 0.0029$
$a_8^{(p)}$	$0.585 \pm 0.025$
$\Delta_8^{(p)}$	$-0.09 \pm 0.03$

Ellis, Olive, Sandick



# ***DM INDIRECT DETECTION***

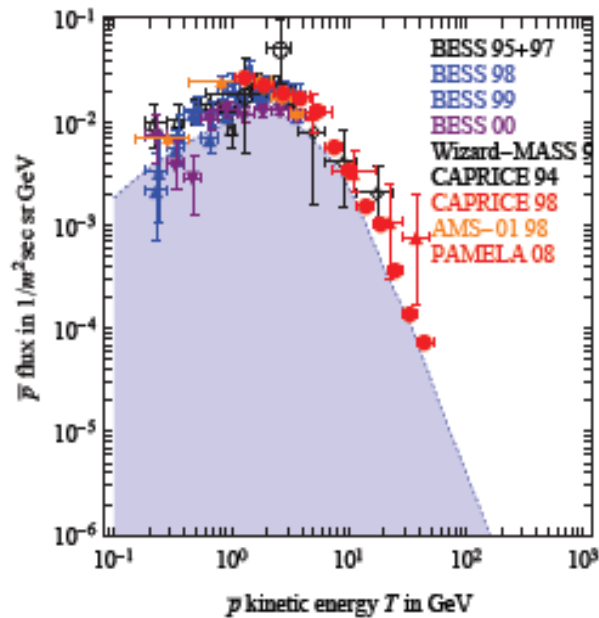
- WIMP-WIMP annihilation in the galactic halos may be detected through production of  $\gamma$ , neutrinos, anti-matter.



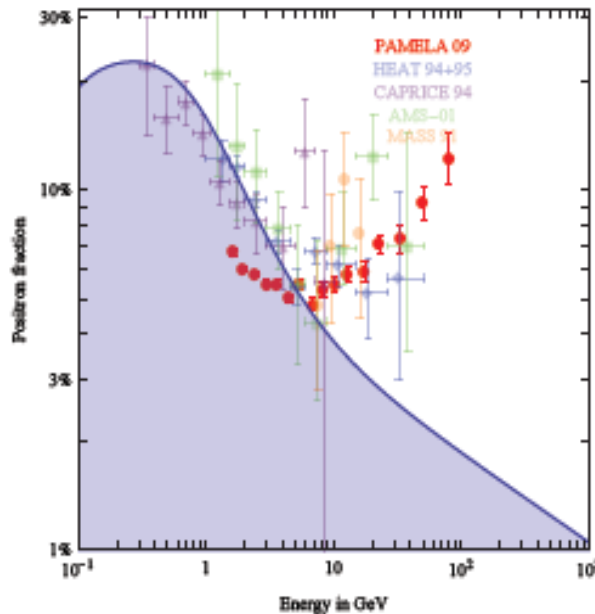
# INDIRECT SEARCHES OF DM

- **WIMPs collected inside celestial bodies** ( Earth, Sun): their annihilations produce energetic neutrinos
- **WIMPs in the DM halo**: WIMP annihilations can take place ( in particular, their rate can be enhanced with there exists a CLUMPY distribution of DM as computer simulations of the DM distribution in the galaxies seem to suggest. From the WIMP annihilation:
  - **energetic neutrinos** ( under-ice, under-water exps Amanda, Antares, Nemo, Nestor, ...)
  - **photons in tens of GeV range** ( gamma astronomy on ground Magic, Hess, ... or in space Agile, Glast...)
  - **antimatter**: look for an excess of antimatter w.r.t. what is expected in cosmic rays ( space exps. Pamela, AMS, ...)

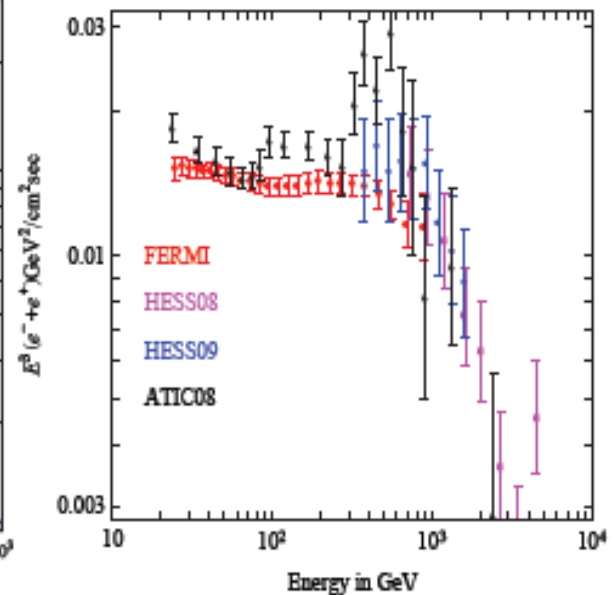
# PAMELA, FERMI/ATIC, HESS



$\bar{p}$ : consistent with bck



$e^+/e^-$ : excess



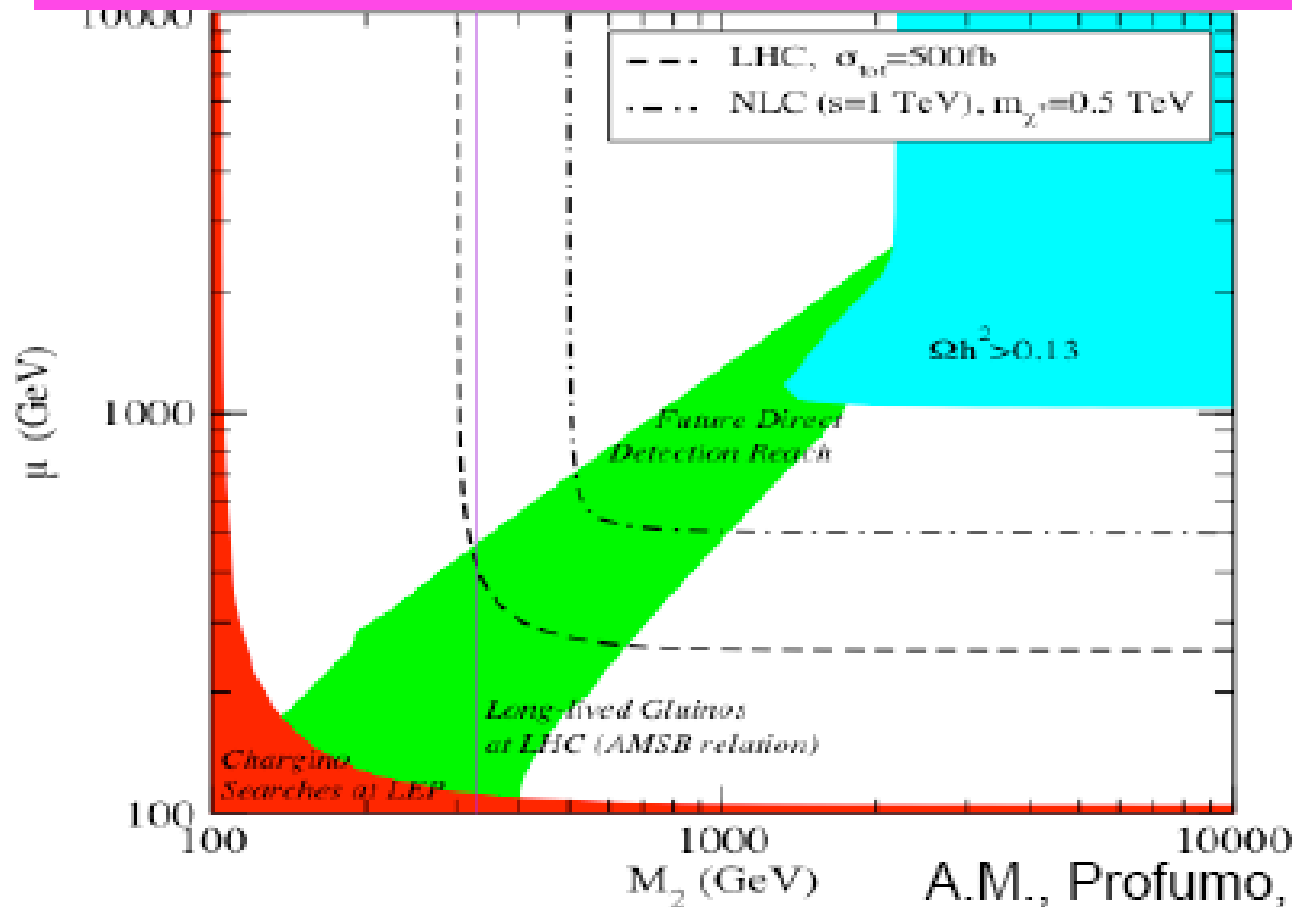
$e^- + e^+$ : feature?

Strumia

EPS09

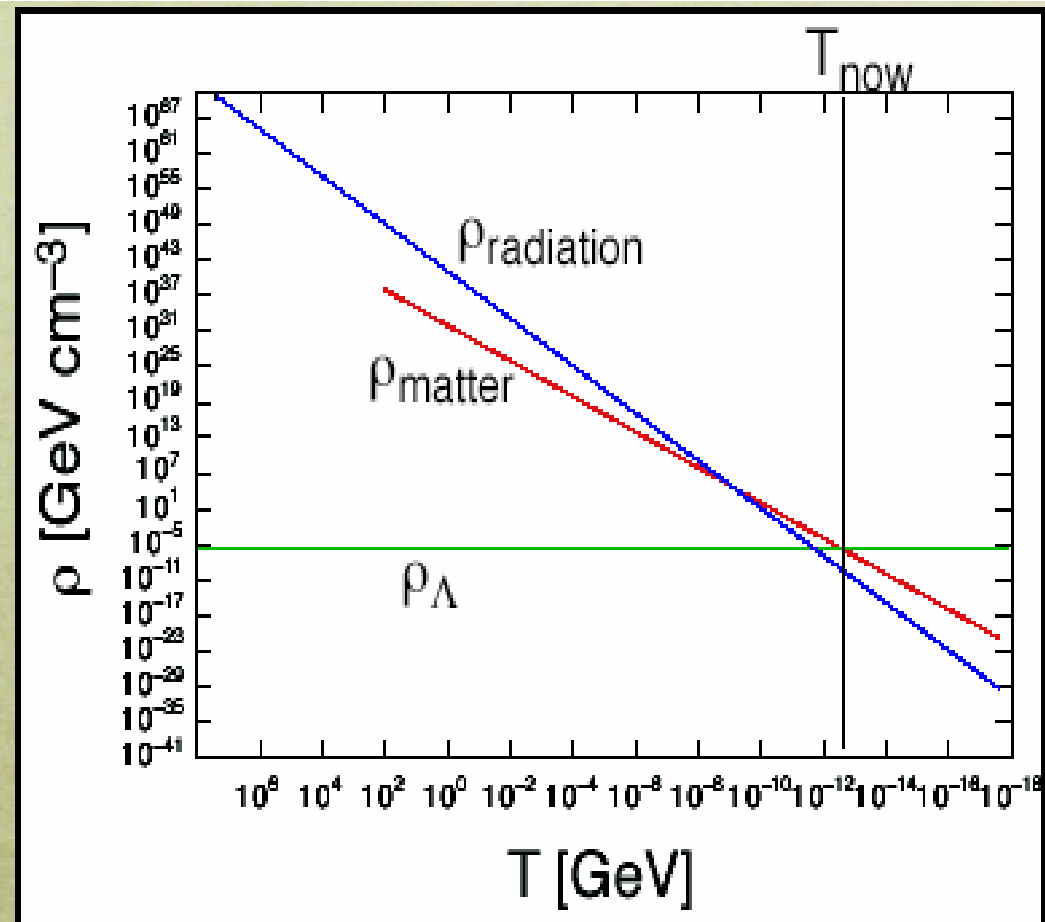
PAMELA excess: October 2008, stimulated enormous theoretical activity; note: statistical errors only! Fermi: feature observed by ATIC not confirmed

# LHC, ILC, DM SEARCHES SENSITIVITIES



# THE “WHY NOW” PROBLEM

- Why do we see matter and cosmological constant almost equal in amount?
- “Why Now” problem
- Actually a *triple coincidence problem* including the radiation
- If there is a deep reason for  $\rho_\Lambda \sim ((\text{TeV})^2/M_{Pl})^4$ , coincidence natural

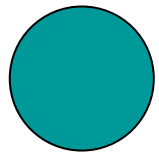


Arkani-Hamed, Hall,  
Kolda, HM

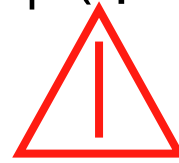




DO THEY "KNOW" EACH OTHER?



DIRECT INTERACTION  $\phi$  (quintessence) WITH DARK MATTER



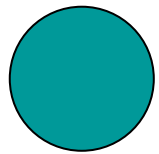
**DANGER:**

$\phi$  Very LIGHT

$m\phi \sim H_0^{-1} \sim 10^{-33} \text{ eV}$



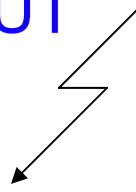
Threat of violation of the equivalence principle  
constancy of the fundamental "constants", ...



**INFLUENCE OF  $\phi$  ON THE NATURE AND THE ABUNDANCE OF CDM**

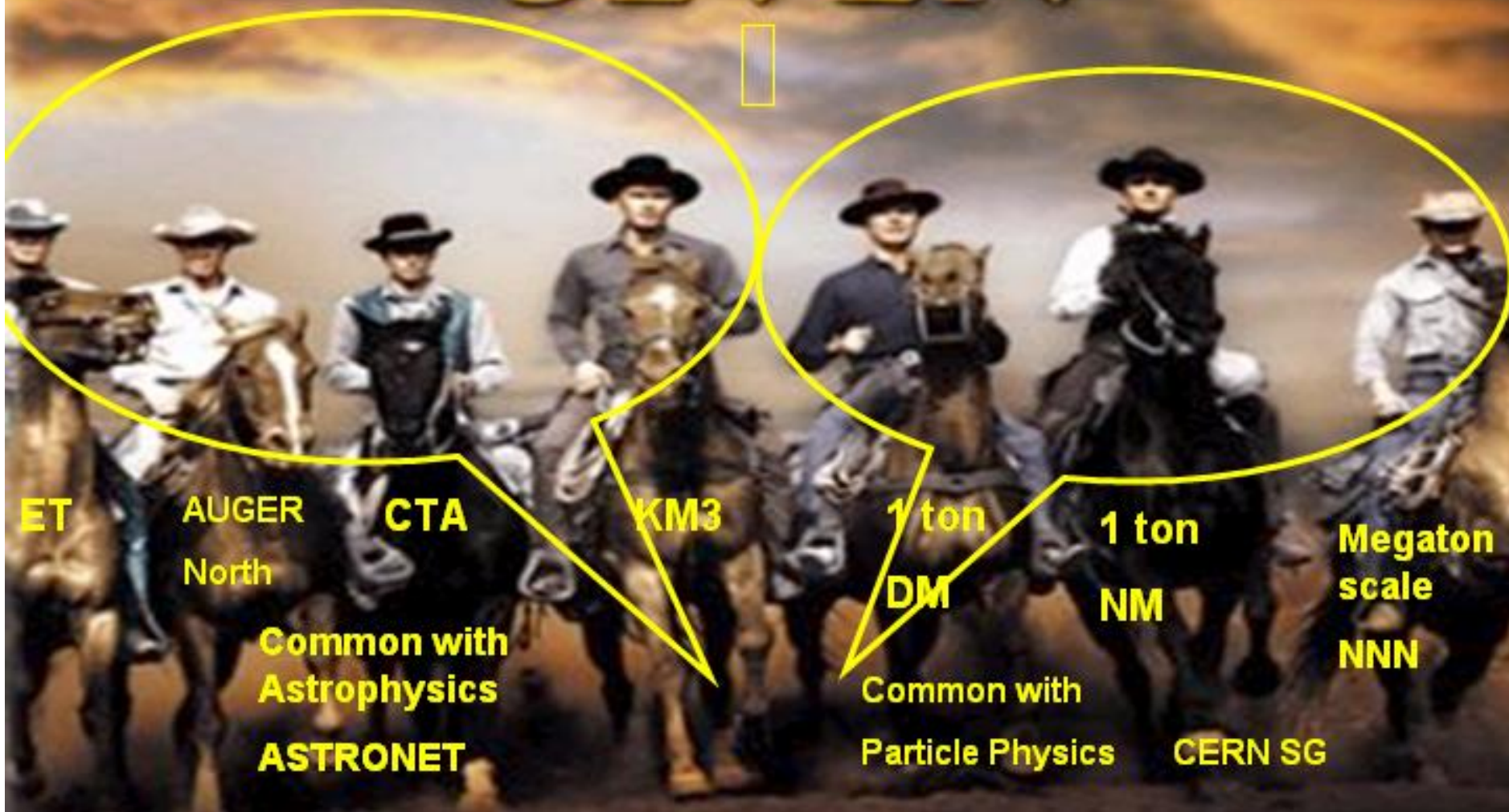
Modifications of the standard picture of  
WIMPs FREEZE - OUT

CDM CANDIDATES



CATENA, FORNENGO, A.M.,  
PIETRONI, SCHELKE

# THE MAGNIFICENT SEVEN



# MICRO

**STANDARD MODEL of  
PARTICLE PHYSICS**

G-W-S MODEL



**BUT ALSO**

# MACRO

**MODELLO STANDARD  
of COSMOLOGY**

HOT BIG BANG



**HAPPY MARRIAGE  
EX: NUCLEOSYNTHESIS**

**FRICION POINTS**

**DARK MATTER AND DARK ENERGY**

**LHC → AN EXCEPTIONAL WINDOW TO  
EXPLORE THE UNIVERSE AND ITS ORIGIN**